



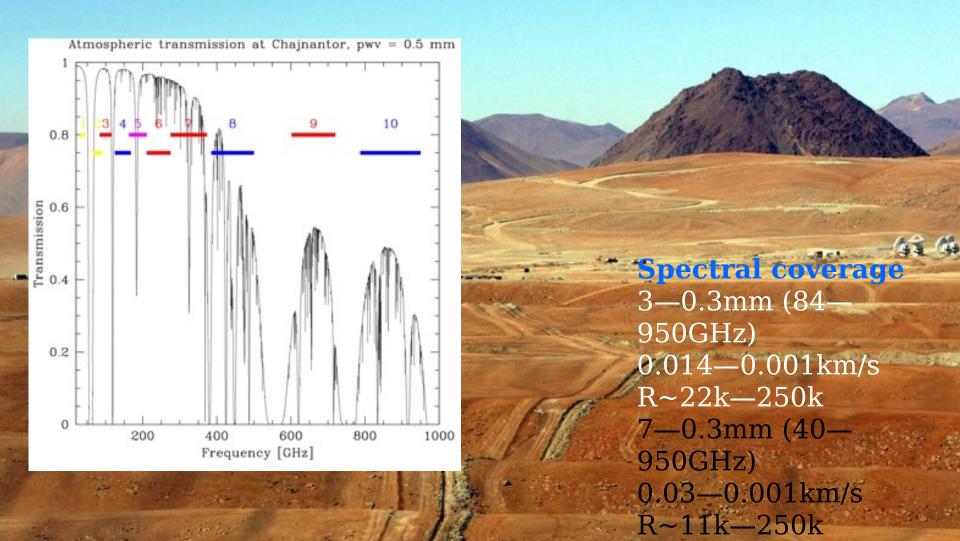
PACDs 2016 14<sup>th</sup> March 2016 www.iastro.pt **Hugo Messias**PACE Lead Scientist hmessias@oal.ul.pt



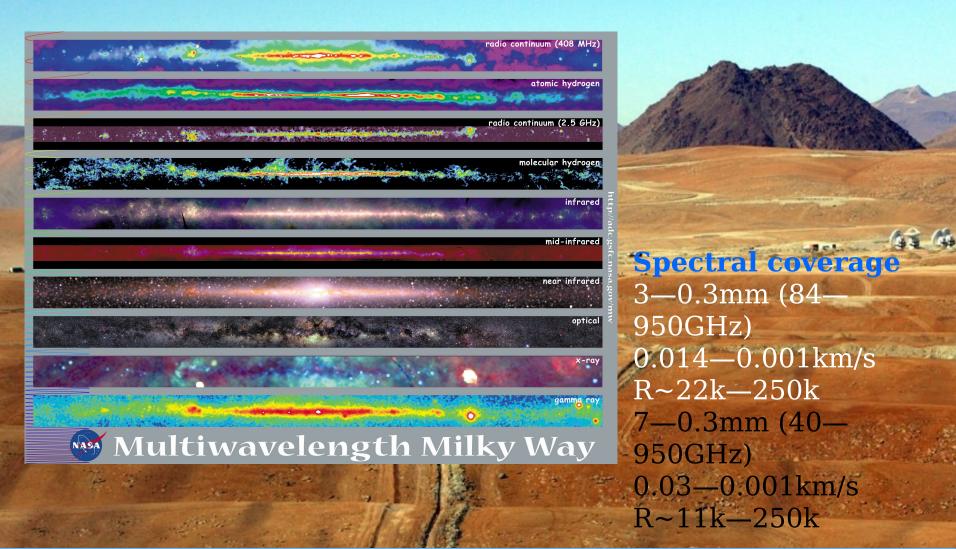








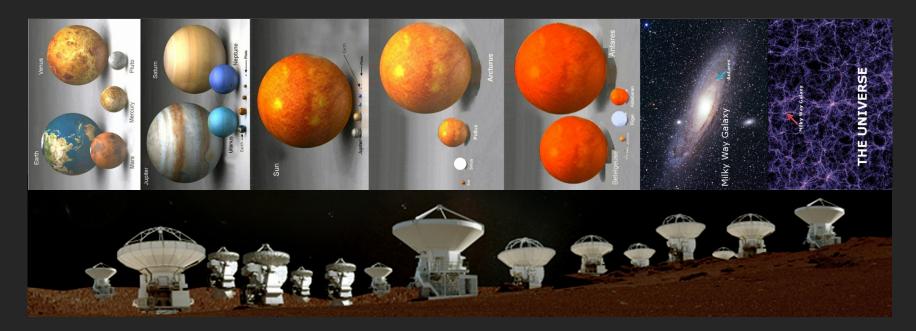






### Science

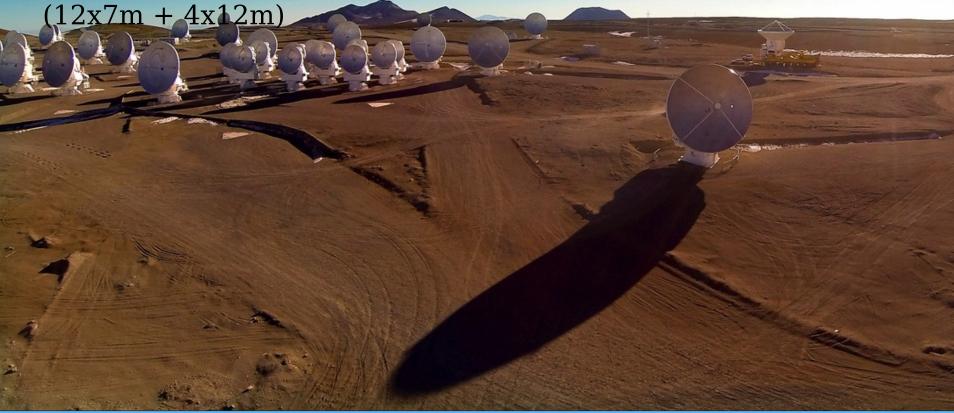
Cosmology and the high redshift universe Galaxies and galactic nuclei ISM, star formation and astrochemistry Circumstellar discs, exoplanets and the solar system Stellar evolution and the Sun



### The array(s)

Main array (12m-array) —  $\geq$ 40 12m antennas (50 in full-ALMA)

Atacama Compact Array (ACA) — ≥10x7m + ≥3x12m



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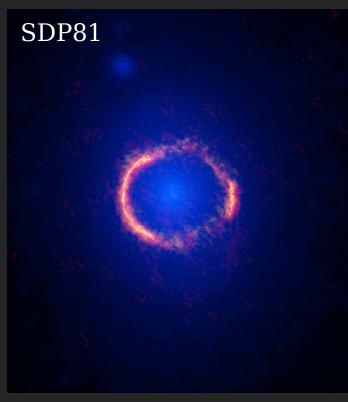
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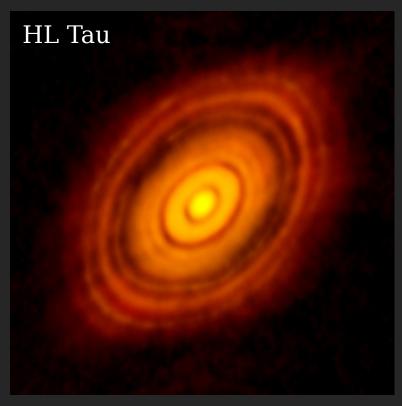


### The array(s)

Main array (12m-array) — <10km apart ( $<\overline{16}$ km)



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© ALMA

### The consortium





### The European ALMA Regional Centre

- distribution of Call for Proposals
- user support for proposal preparation
- TAC procedures and technical feasibility
- assistance with phase II scheduling/execution observations
- data products support
- archive operations
- ALMA helpdesk
- community development and outreach

### The European ALMA support network



Portuguese
ALMA
Centre of
Expertise



PACE

Hatziminaoglou et al. (2015, Msngr, 162

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- user support for proposal preparation
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### **PACE**

The above, except \*, and:

- develop unique expertise (e.g., planetary atmospheres, data mining)
- increase ALMA community usage (e.g., PACDs, talks, ...)
- EU ARC activities (e.g., meetings)

### PACE — team

Coordination — José Afonso

Lead Scientist — Hugo Messias

Astronomers — Ciro Pappalardo, Silvio Lorenzoni, Israel

Matute, Pedro Machado, Sonia Antón, Luca Bizzocchi

Technical Support — Carlos Santos

Administrative Support — Sandra Homem

Public outreach — João Retrê





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#### **PACE NEWS**

ALMA Cycle 4 preannouncement

ALMA Fellowships

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Water in the Universe – from clouds to oceans

ALMA Cycle-3 outcome

#### PACE MAILING LIST

Do you wish to receive ALMArelated news or notices of events organised by the EU ARC or PACE?

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### PACE

Since May 2014, the Institute for Astrophysics and Space Science (Lisbon node, IA) is officially part of the European Atacama Large Millimetre Array (ALMA) support structure as a Centre of Expertise (CoE). This status was granted by ESO after the recognition of IA team's capability to support the community with the use of ALMA, in addition to the already existent EU ALMA Regional Centre (ARC) nodes.

The Portuguese ALMA CoE (PACE) is now composed by a <u>team</u> of researchers, technical, administrative, and outreach personnel. <u>IA</u> members are currently involved in seven approved ALMA proposals. The <u>tasks</u> of the PACE are partly the same as those of an EU ARC node, including, for instance, proposal preparation support.

http://pace.oal.ul.pt



PACDs 2016 14<sup>th</sup> March 2016

The observatory ALMA
Hugo Messias (PACE)



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### ALMA Cycle 4 Pre-announcement

Dec 14, 2015

The Joint ALMA Observatory (JAO) will start the next cycle of observing (Cycle 4) in October 2016. A Call for Proposals with detailed information on Cycle 4 will be issued in March 2016, with a deadline for proposal submission in April 2016. This pre-announcement highlights aspects of the Cycle 4 proposal call that are needed to plan proposals.

#### General information

ALMA Cycle 4 will start in October 2016 and span 12 months. It is anticipated that 3000 hours of 12-m Array science observations will be available for successful observations of approved projects. Observing time will also be available on the Morita Array (a.k.a. the Atacama Compact Array, or the ACA); the number of hours available on the ACA will be announced in the Call for Proposals. The remaining time on ALMA will be reserved for engineering, computing and scientific testing to extend and optimize ALMA capabilities.

The key dates for Cycle 4 are given below.

22 March 2016	Release of the ALMA Cycle 4 Call for Proposals and observing tool, and opening of archive for proposal submission
21 April 2016	Proposal deadline
August 2016	Result of the proposal review process sent to Proposers
October 2016	Start of ALMA Cycle 4 observations
September 2017	End of Cycle 4 observations

## http://almascience.eso.org/



### Cycle 4 — anticipated capabilities

- Number of antennas:  $\geq$ 40 12m-array; 10 7m-array; 3 12m
- Bands: 3, 4, 6, 7, 8, 9, 10 (3.1, 2.1, 1.3, 0.87, 0.74, 0.44,
- 0.35mm, respectively)
- 12m-array configurations:
  - Nine with maximum distances of 155m to 12.6km
  - Maximum 2.7km for Bands 8, 9, 10
  - Maximum 5.3km for Band 7
  - Maximum 12.6km for Bands 3, 4, 6

### Cycle 4 — anticipated capabilities

- All bands: spectral line and continuum single field observations with 12m and 7m arrays
- Bands 3 to 9: mosaics with 12m and 7m arrays
- Bands 3 to 8: single dish spectral line observations
- Bands 3, 6, 7: single pointing, on axis, full (linear) polarization for continuum and full spectral resolution observations with 12m array

### Cycle 4 — non-standard modes

- Bands 8, 9, 10
- Band 7 with maximum baselines >2.7km
- Polarization
- Spectral scans
- <1GHz aggregate bandwidth over all spectral windows
- Solar
- VLBI
- User-specified calibrations

### Cycle 4 — new opportunities

- ACA stand-alone mode
- Large programmes (>50h)
- VLBI (Bands 3 and 6)
- Solar observations (Bands 3 and 6)





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